

Freeing semiclassical theory of the dreaded UV divergence



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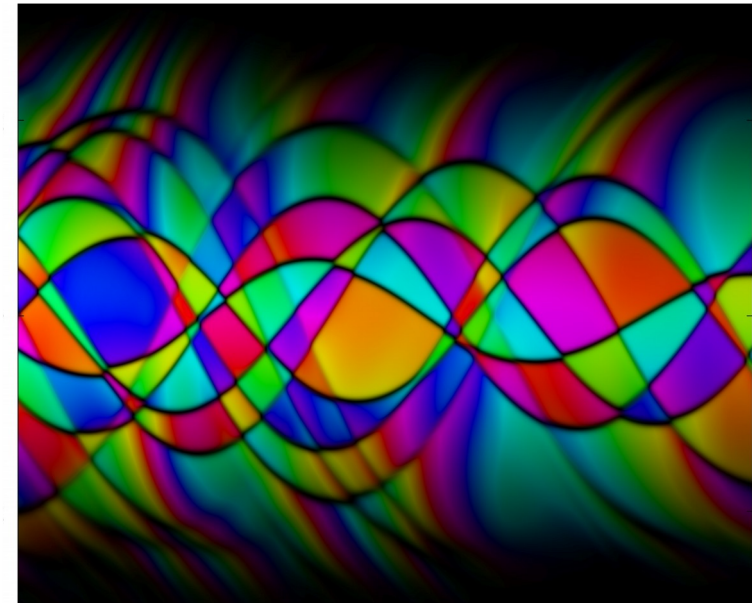


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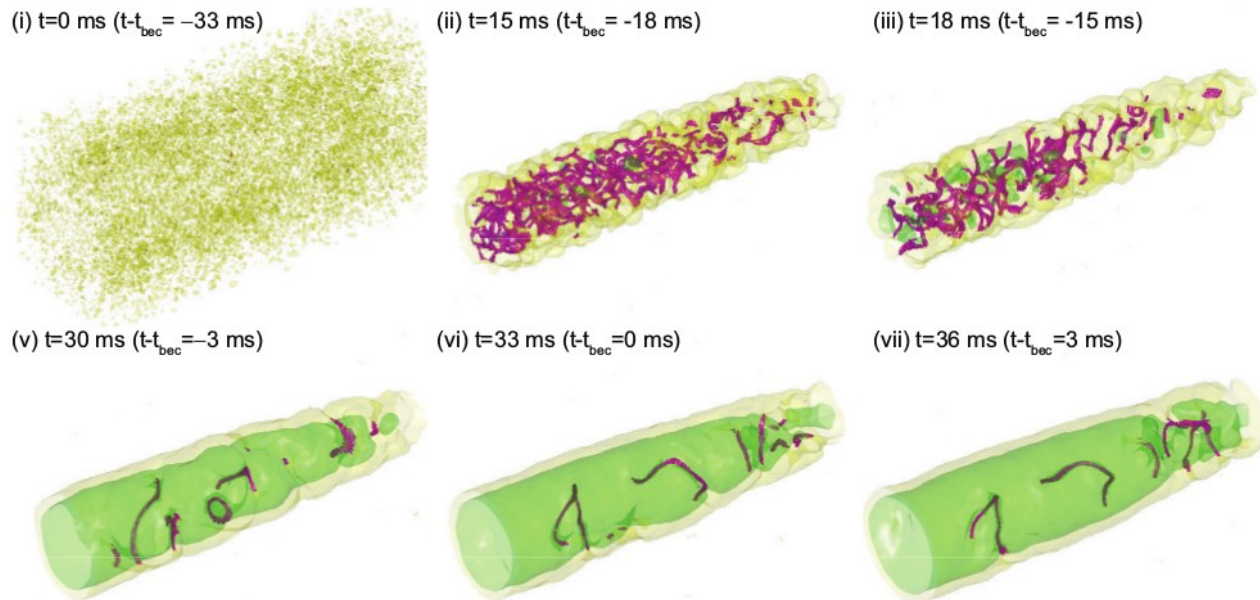
- Semiclassical wave fields for bosons (background)
- The UV divergence and cutoff issues
- Fixing it
- Simple example: 1d
(~1000 modes)
- Serious example: collective breathing at high temperature in 3d
(1000000s of modes)

The point of semiclassical field theory

- Nonperturbative
- $T > 0$
- Single shots
- No special conditions, symmetries, geometries
- Scales very well with system size



1d gas – phase domains after cooling

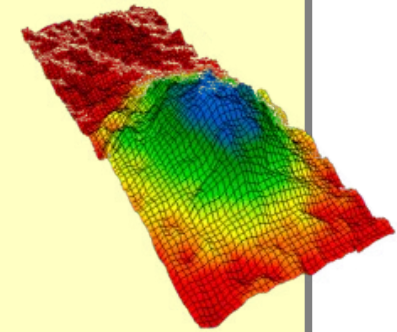
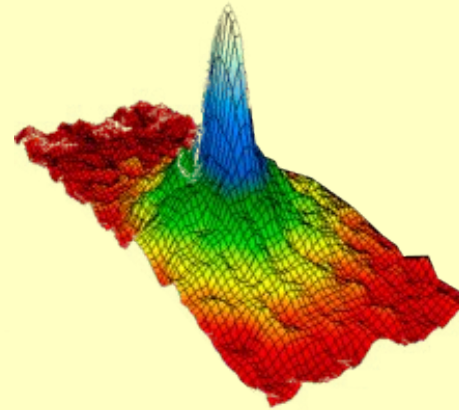
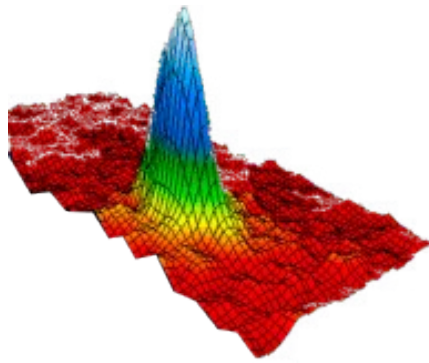


$$\langle n(x, t) \rangle$$

$$n(x, t)$$

1d gas – thermal equilibrium

Thermal clouds - Classical fields essential



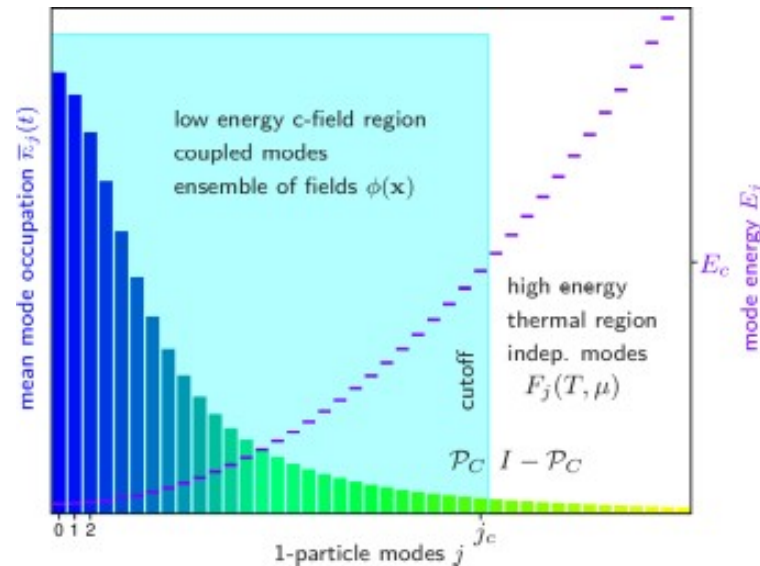
$T=0$
1 mode
GPE

Bogoliubov
few modes
perturbative

Classical fields, ZNG
zillions of modes
non-perturbative

$T=T_c$

Simplest semiclassical method = “classical” wave fields



Assuming high occupation:

Bose field semiclassical replacement

$$\hat{\Psi}(\mathbf{x}) = \sum_j \hat{a}_j \psi_j(\mathbf{x}) \longrightarrow \phi(\mathbf{x}) = \left\{ \sum_{j \in \mathcal{C}} \alpha_j \psi_j(\mathbf{x}) \right\}$$

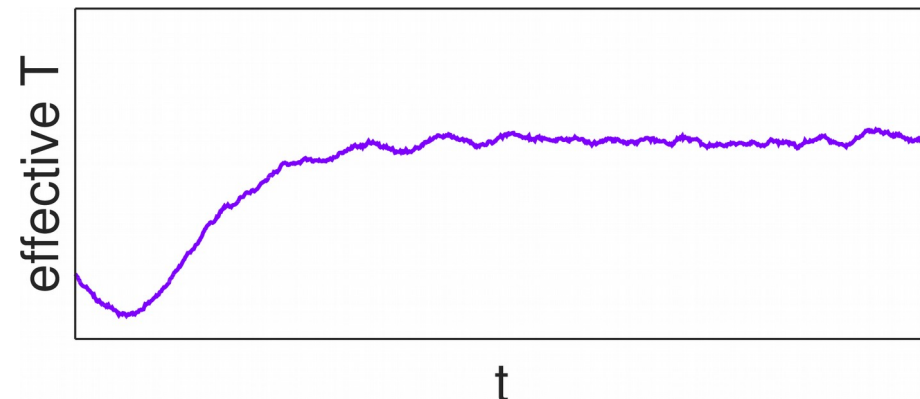
$$\begin{aligned} [\hat{a}_j, \hat{a}_k^\dagger] &= \delta_{jk} \\ \hat{a}_j \gg 1 &\rightarrow \hat{a}_j \approx \alpha_j \end{aligned}$$

Evolution: GPE

$$\hbar \frac{d\phi(\mathbf{x})}{dt} = -i\mathcal{E}(\mathbf{x})\phi(\mathbf{x})$$

$$\mathcal{E}\phi(\mathbf{x}) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\mathbf{x}) + g|\phi(\mathbf{x})|^2 \right] \phi(\mathbf{x})$$

Thermalised ensemble at long time



Methods for thermal bose gases

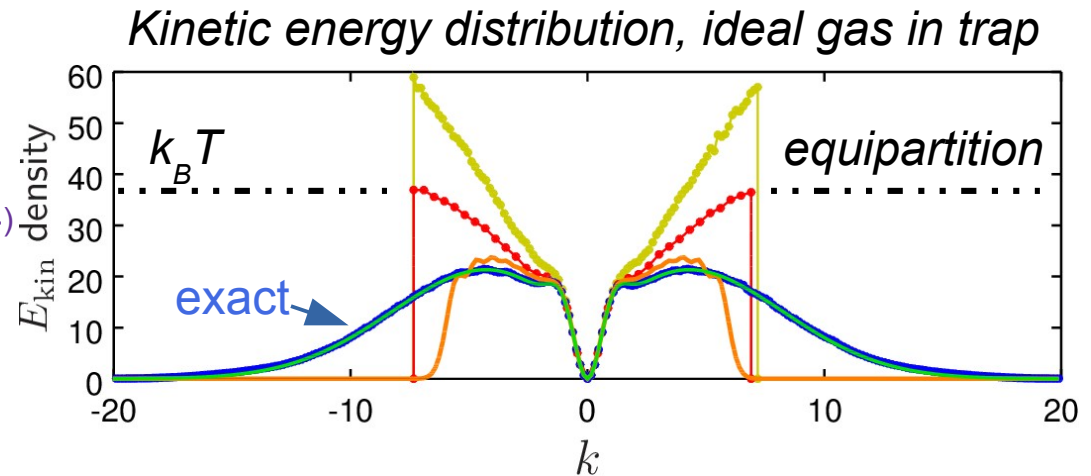
	ZNG	GPE (classical fields)	Projected GPE + Hartree-Fock	Stochastic GPE
nonperturbative low E modes	✓	✓	✓	✓
many modes at low E	✗	✓	✓	✓
High E modes	dynamic	✗	static	reservoir
temperature	?	extracted	extracted	set
Equilibrium ensemble	?	micro-canonical	micro-canonical	grand-canonical
UV divergence Cutoff dependence	no	STRONG	some	STRONG
numerical effort	LARGE	medium	medium	medium

WANTED: a way to combine the good features of ZNG and classical fields

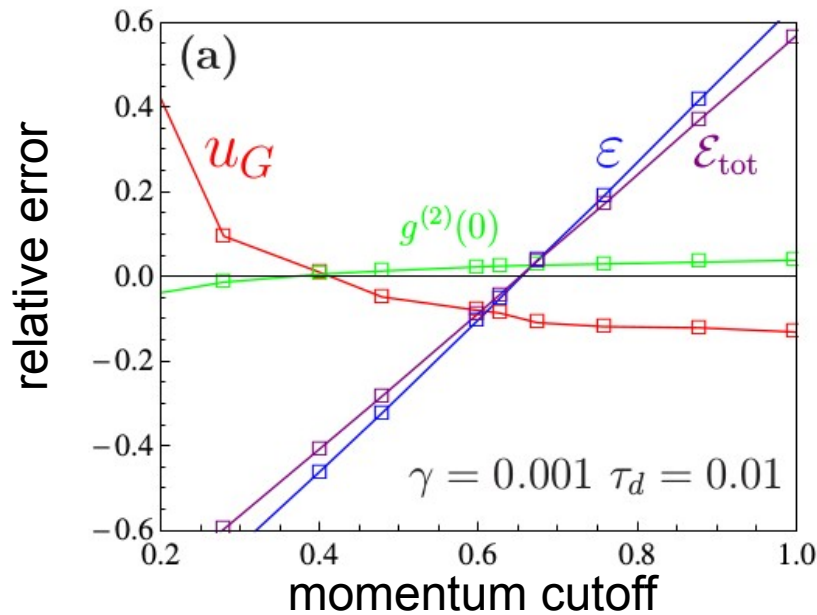
Age-old problem: cutoff dependence

Studied by many:

- Witkowska, Gajda, Rzażewski, PRA **79**, 033631 (2009)
 - Karpiuk, Brewczyk, Gajda, Rzażewski, PRA **81**, 013629 (2010)
 - Zawitkowski, Brewczyk, Gajda, Rzażewski, PRA **70**, 033614 (2004)
 - Bradley, Blakie, Gardiner, J Phys B **38**, 4259 (2005)
 - Cockburn, Proukakis, PRA **86**, 033610 (2010)
- and the list goes on ...*

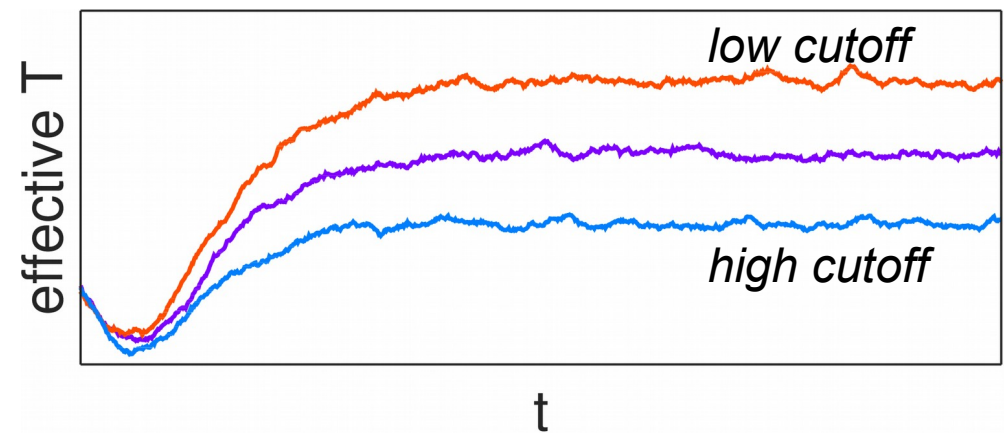


Most recent study: different observables



- Pietraszewicz, PD, PRA **92**, 063620 (2015)
- Pietraszewicz, PD, PRA **97**, 053607 (2018)
- Pietraszewicz, PD, PRA **98**, 023622 (2018)

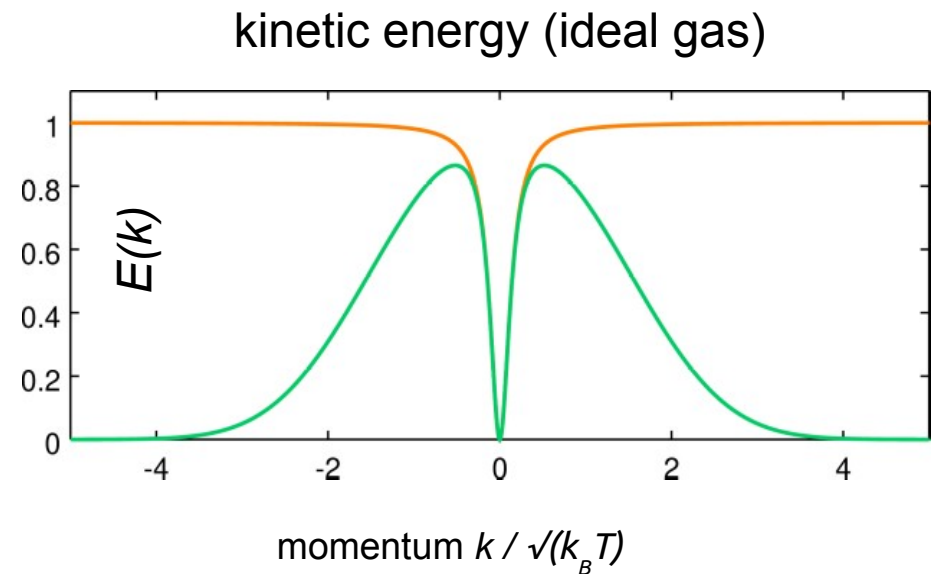
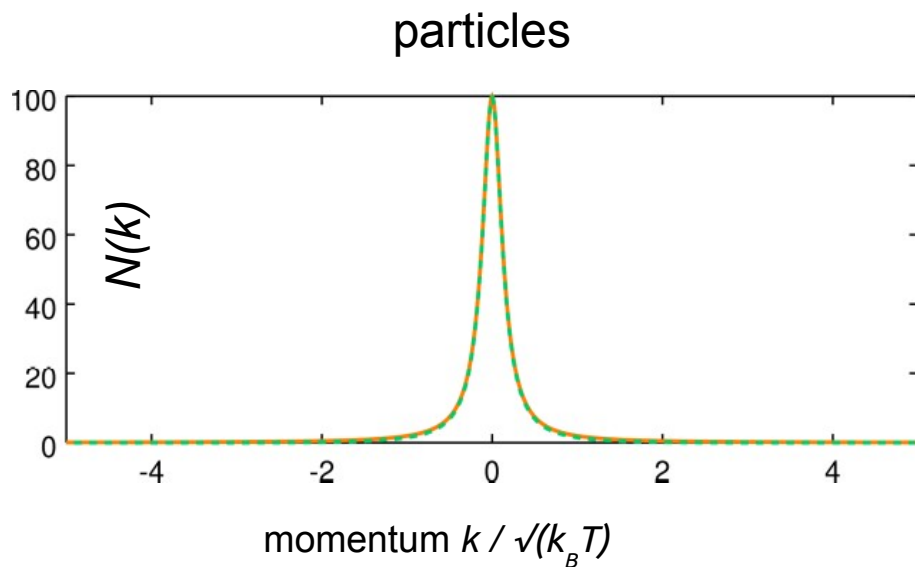
Equilibrium temperature depends on cutoff



Ultraviolet catastrophe = nefarious equipartition

Quantum Bose-Einstein distribution $\frac{1}{2}k_B T$ energy per *PARTICLE*

Classical field Rayleigh-Jeans distribution $k_B T$ energy per *MODE*



Also applies to truncated Wigner at long times

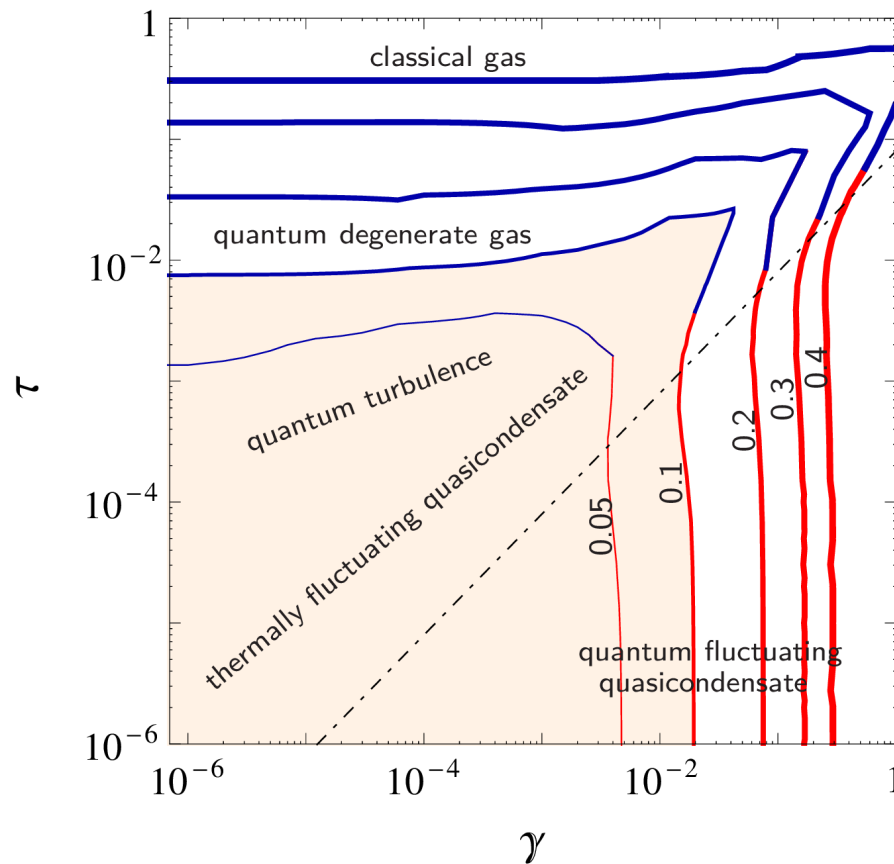
- Uniform 1d Bose gas: comparison to exact results

$$\gamma = \frac{mg}{\hbar^2 n}; \quad \tau_d = \frac{T}{T_d}$$

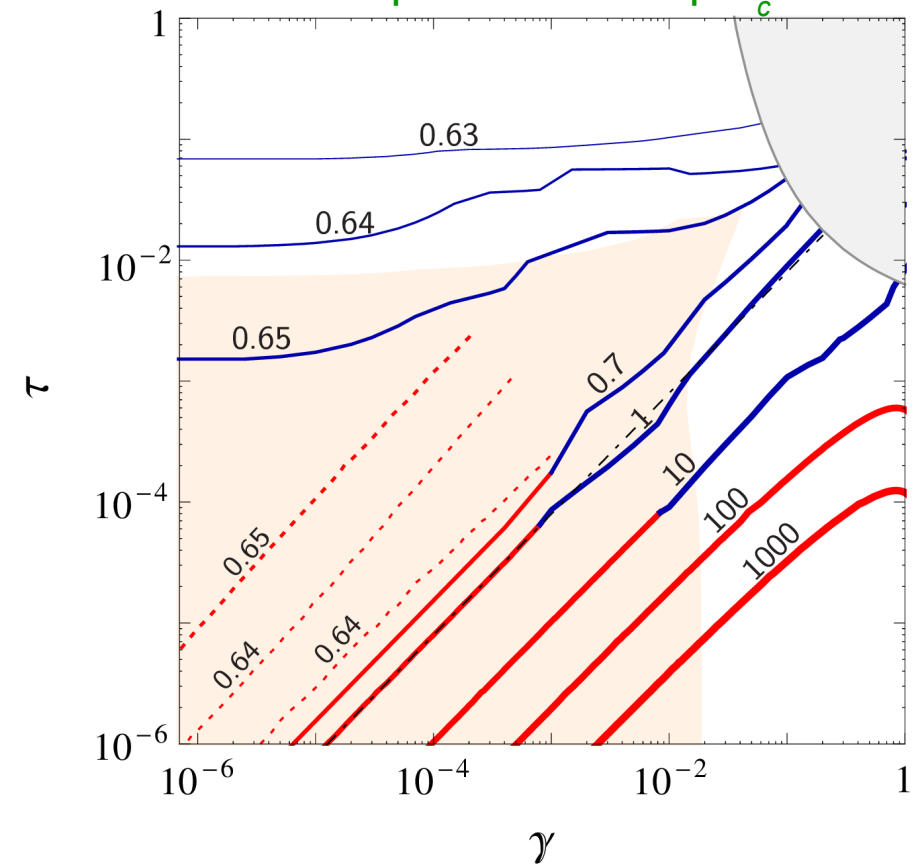
Relative error $\delta_\Omega(\gamma, \tau_d, f_c) := \frac{\Delta\Omega}{\Omega} = \left(\frac{\Omega^{(cf)}(\gamma, \tau_d, f_c)}{\Omega(\gamma, \tau_d)} - 1 \right)$

Global error bound $RMS(\gamma, \tau_d, f_c) = \sqrt{\left(\delta_{u_G} \right)^2 + \max\left[\delta_\varepsilon^2, \delta_{\mathcal{E}_{tot}}^2 \right]}$

Global accuracy of observables min RMS

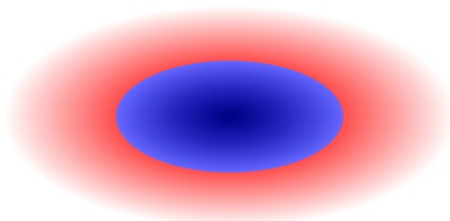


Optimal cutoff $opt f_c$

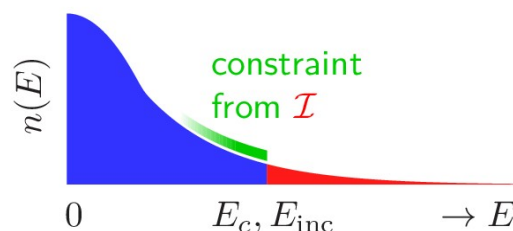


Stochastic Gross-Pitaevskii equation (SGPE)

real space



occupations



complex noise

$$\hbar \frac{\partial \phi(\mathbf{x})}{\partial t} = -i\mathcal{E}(\mathbf{x})\phi(\mathbf{x}) - \gamma [\mathcal{E}(\mathbf{x}) - \mu] \phi(\mathbf{x}) + \sqrt{2\hbar\gamma k_B T} \eta(\mathbf{x}, t)$$

Hamiltonian evolution of
"coherent" field

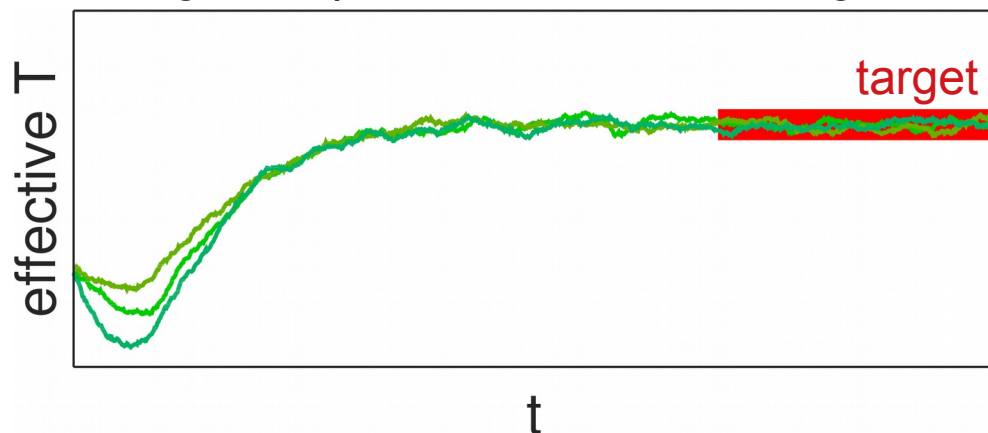
Loss rate to "incoherent" tails

Incoherent growth from tails

$$\mathcal{E}\phi(\mathbf{x}) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\mathbf{x}) + g|\phi(\mathbf{x})|^2 \right] \phi(\mathbf{x})$$

BUT, still UV divergent

Target temperature reached at long time



Invokes a "classical field"
linearisation of occupation in tails
(correct when $\bar{N}(\omega) \gg 1$)

$$\bar{N}(\omega) \rightarrow \frac{k_B T}{\omega(\mathbf{x}) - \mu}$$

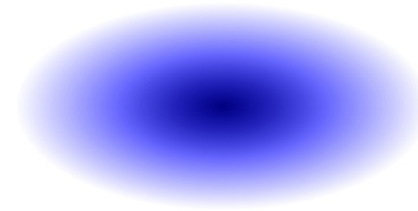
→ equipartition strikes again :(

Idea: do the distribution in the tails properly, already

Use proper quantum occupations

$$\bar{N}(\omega) = N_{BE} = \left[e^{\frac{\omega(\mathbf{x}) - \mu}{k_B T}} - 1 \right]^{-1}$$

real space



occupations



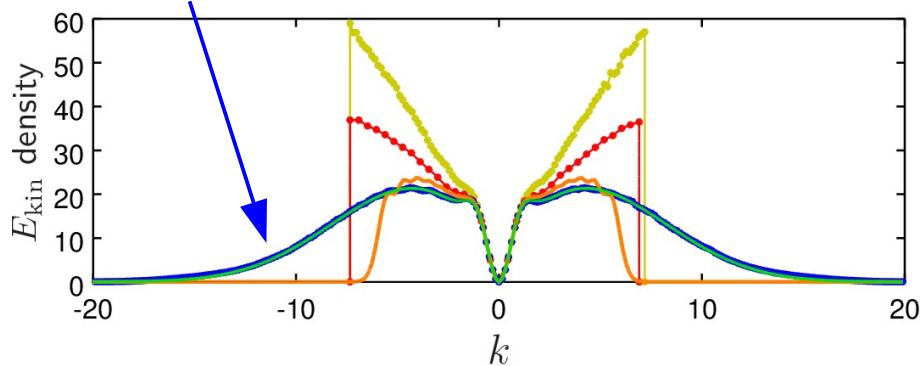
Obtain regularised SGPE

$$\hbar \frac{\partial \phi(\mathbf{x})}{\partial t} = -i\mathcal{E}(\mathbf{x})\phi(\mathbf{x}) - \gamma k_B T \left[e^{\frac{\mathcal{E}(\mathbf{x}) - \mu}{k_B T}} - 1 \right] \phi(\mathbf{x}) + \sqrt{2\hbar\gamma k_B T} \eta(\mathbf{x}, t)$$

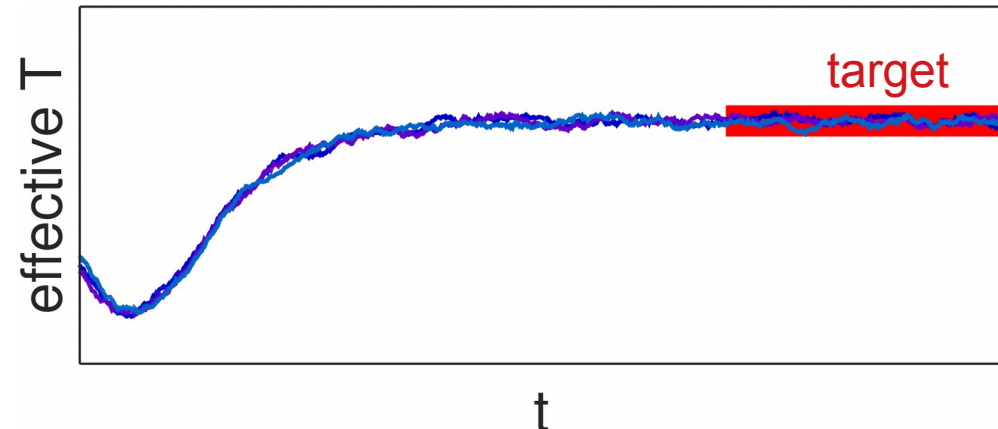
$$\mathcal{E}\phi(\mathbf{x}) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\mathbf{x}) + g |\phi(\mathbf{x})|^2 \right] \phi(\mathbf{x})$$

“Only” real difference

UV divergence is gone

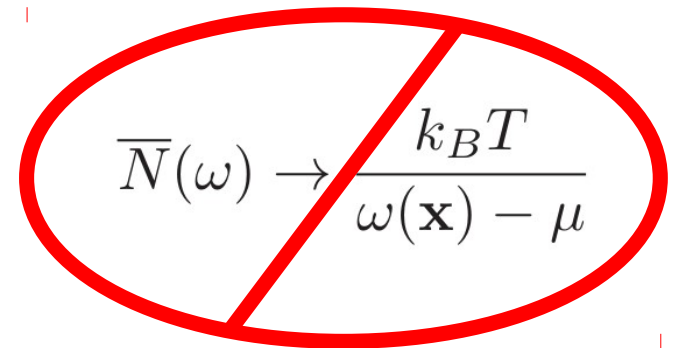


Evolution well controlled



Derivation (rather abridged)

- [Start with master equation for stochastic GPE model](#)
 - * **BUT DO NOT linearize bath**
- Keep only one-particle exchange terms with reservoir
 - * Like in SGPE derivation
- Reduce reservoir coupling to a localised form
 - * Happened automatically in SGPE, need to take care here
- Represent master equation using positive-P distribution
 - * Needed to pinpoint how to properly remove quantum noise (problem didn't occur in SGPE)
- Convert to Fokker-Planck equation
- Convert to stochastic equation
- Remove quantum noise terms which cause instability (noise amplification)
- Obtain regularised SGPE:

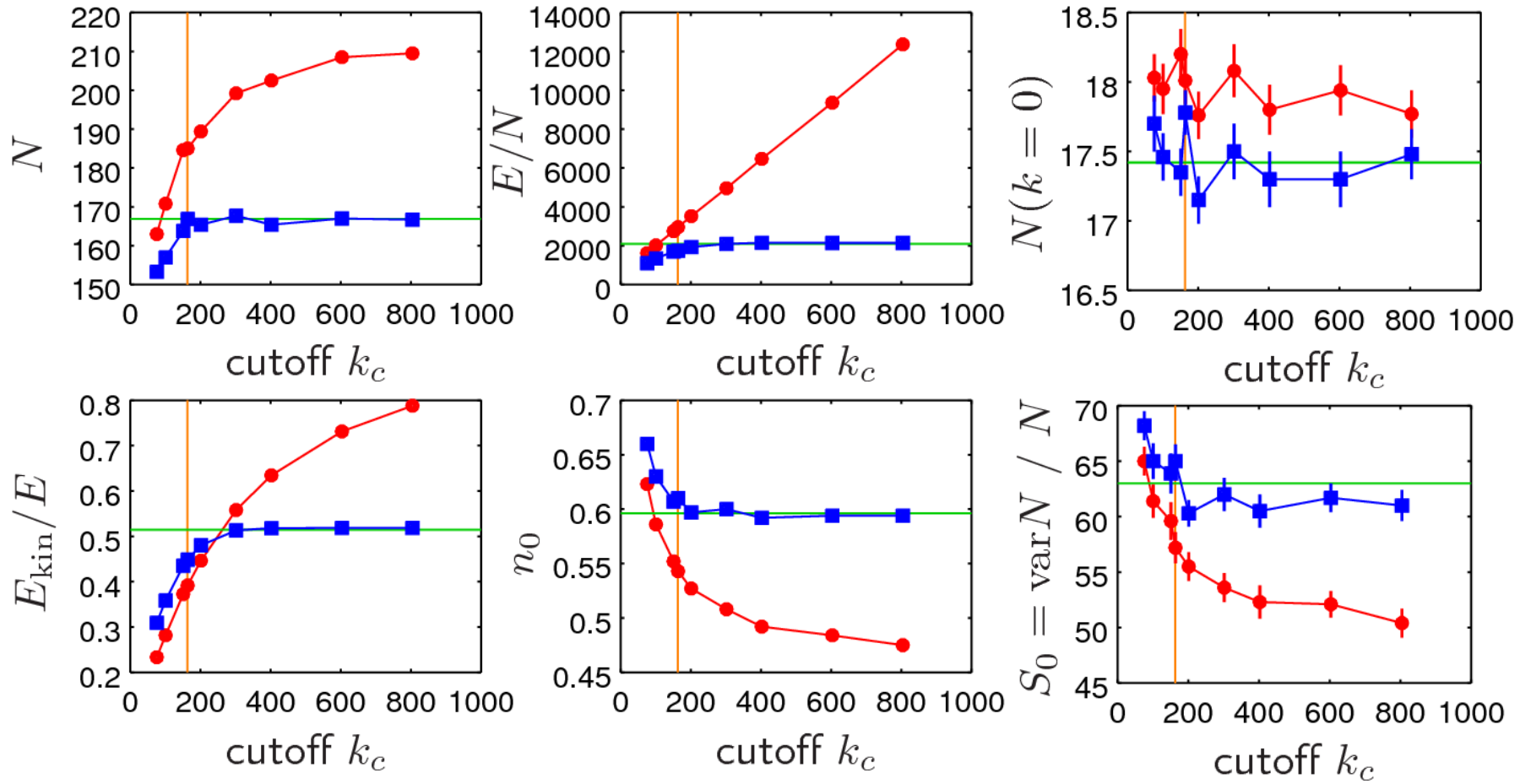

$$\bar{N}(\omega) \rightarrow \frac{k_B T}{\omega(\mathbf{x}) - \mu}$$

Stubborn term. Not diagonal; and likes to be huge.

$$\hbar \frac{\partial \phi(\mathbf{x})}{\partial t} = -i\mathcal{E}(\mathbf{x})\phi(\mathbf{x}) - \gamma k_B T \left[e^{\frac{\mathcal{E} - \mu}{k_B T}} - 1 \right] \phi(\mathbf{x}) + \sqrt{2\hbar\gamma k_B T} \eta(\mathbf{x}, t)$$

- [Implement timestep in a special way](#) to preserve the tractable computational cost of plain SGPE
 - * Was not needed in SGPE, there was no exponential.
 - * This is tricky. Probably the reason the equation was not tried earlier.

Ideal 1d trapped gas (cutoff dependence)



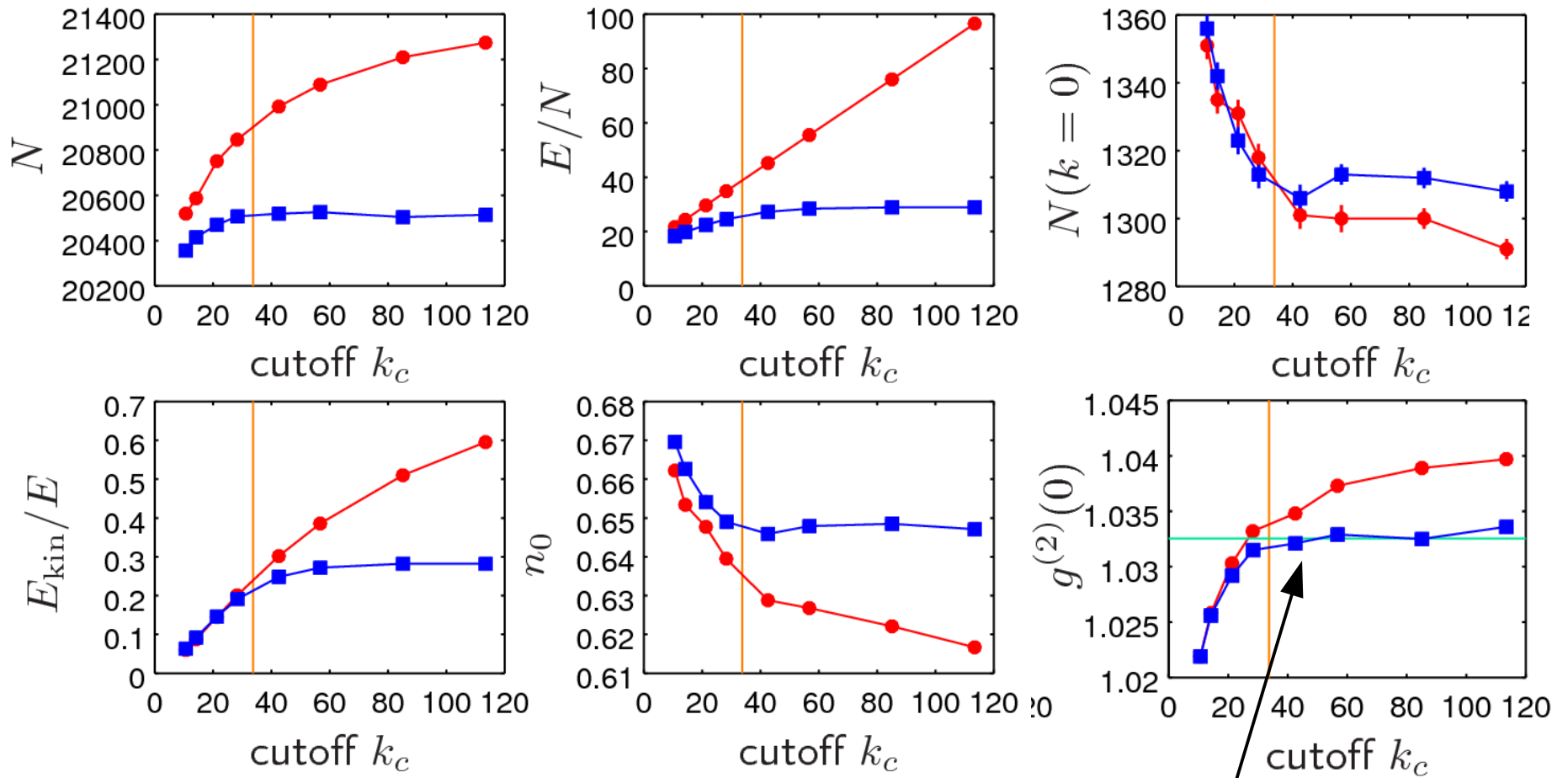
exact

plain SGPE

modified SGPE

cutoff from [arXiv:1707.01776](https://arxiv.org/abs/1707.01776)

Interacting 1d trapped gas



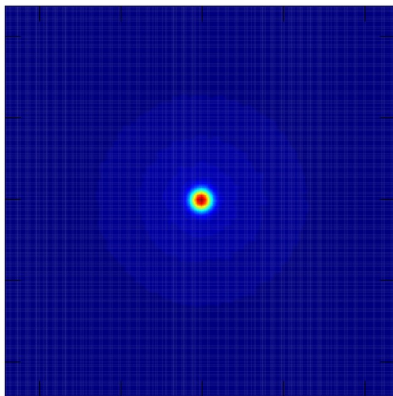
plain SGPE (classical field)

modified SGPE

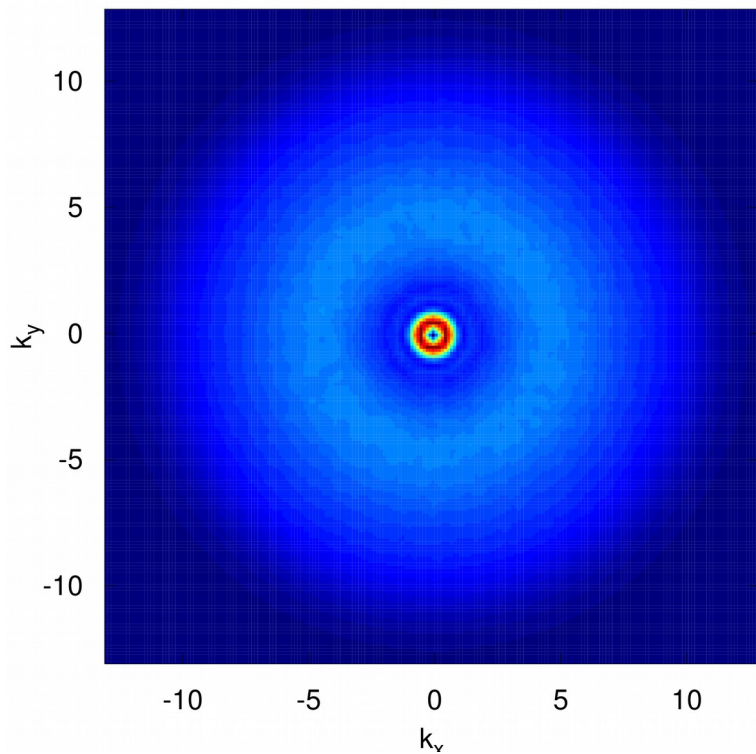
cutoff from arXiv:1707.01776

extended Bogoliubov as per Mora+Castin PRA 67, 053615 (2003)

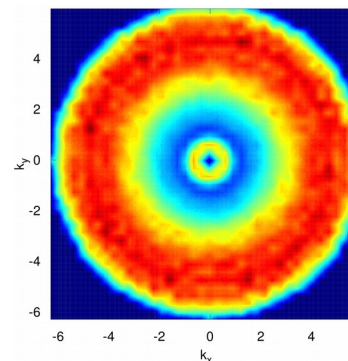
density in momentum space



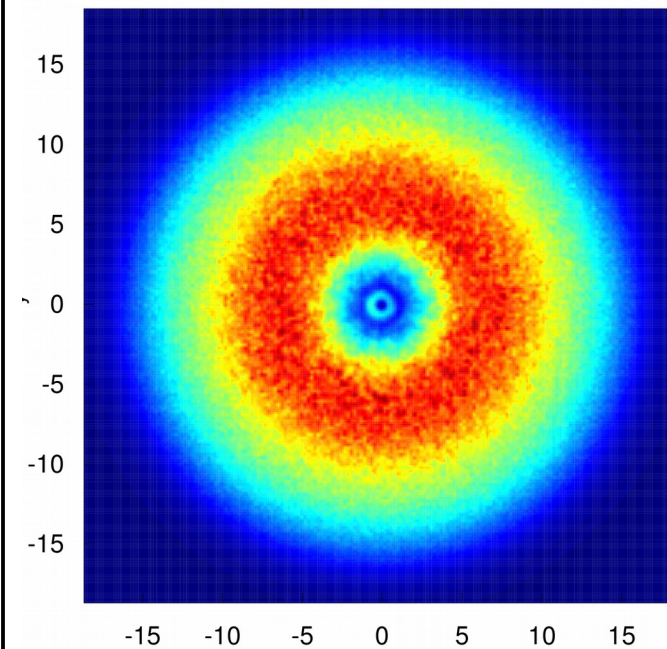
kinetic energy density



plain SGPE – to scale

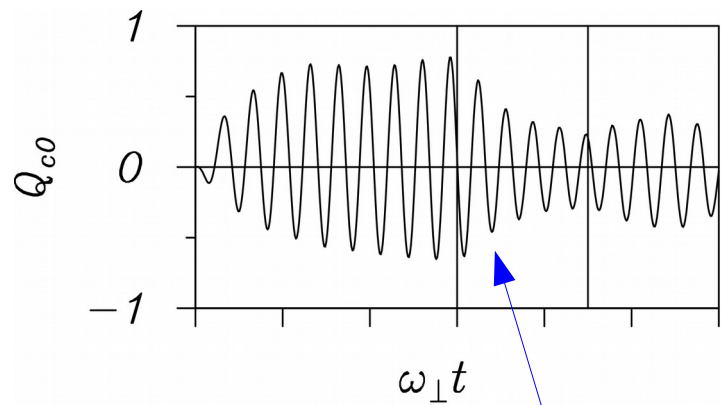
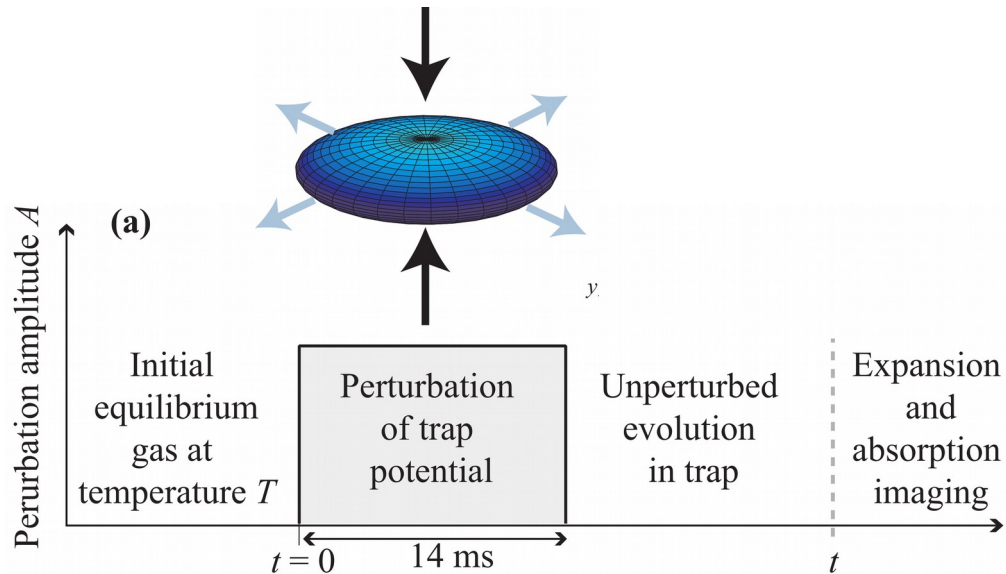


higher temperature

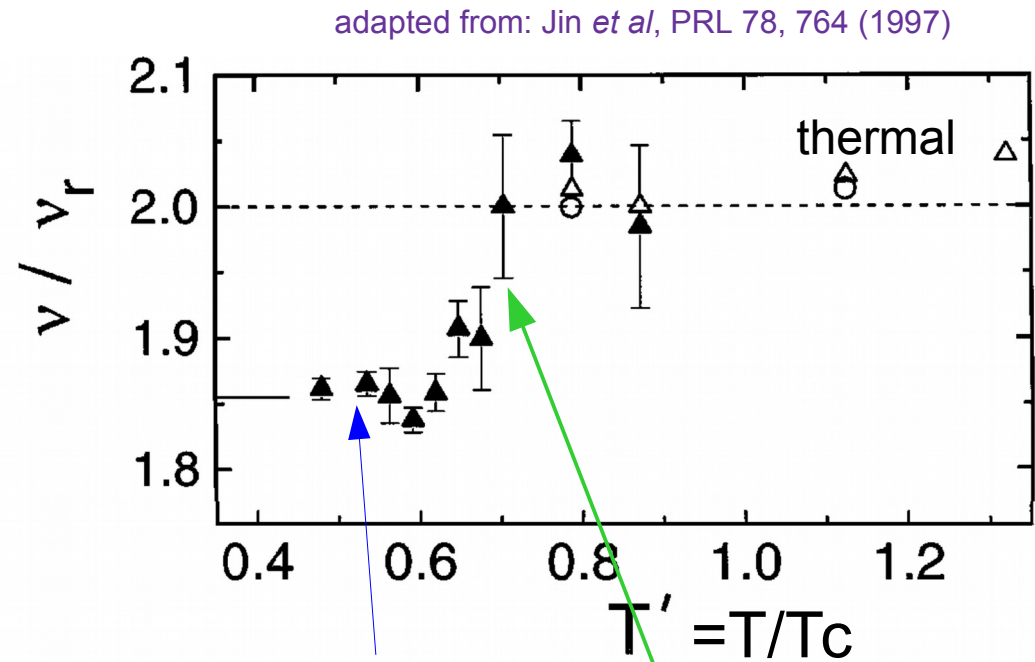


Litmus test for thermal cloud: $m=0$ collective mode

Ancient evil experiment: Jin, Matthews, Ensher, Wieman, Cornell, PRL 78, 764 (1997)

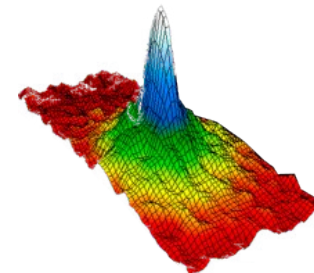


fit decaying sinusoid



condensate oscillations

shift due to drag of thermal cloud on condensate



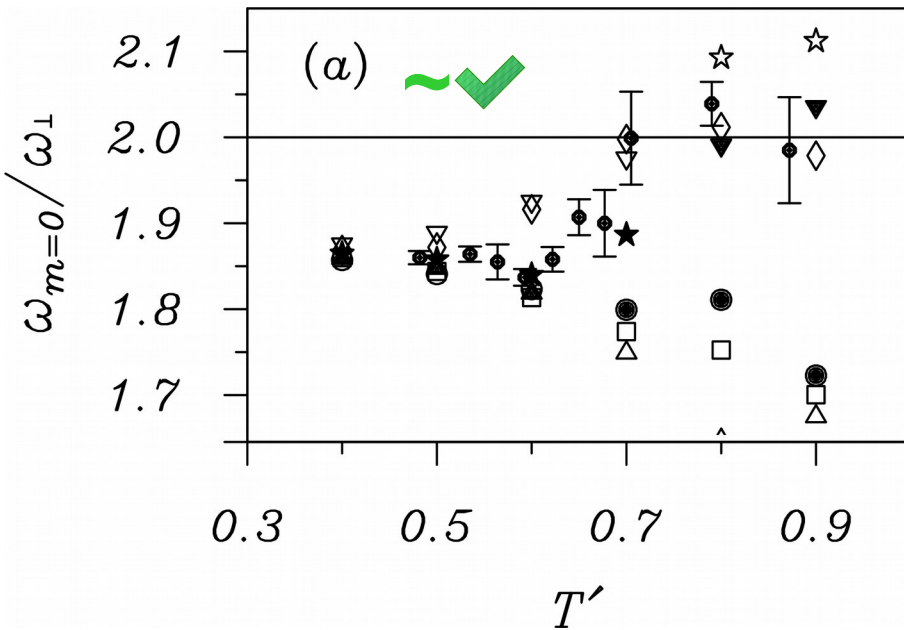
The above figures from:
 Jackson, Zaremba, PRL 88, 180402 (2002)
 Bezett, Blakie, PRA 79, 023602 (2009)

Results of $m=0$ mode tests in the past

Targeted approaches

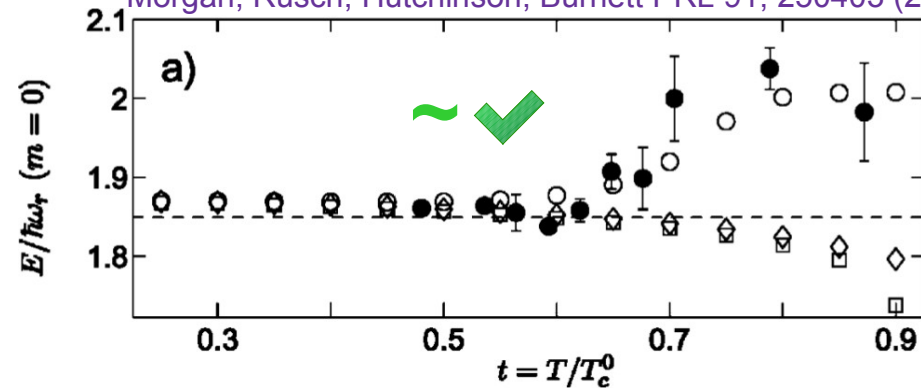
ZNG

Jackson, Zaremba, PRL 88, 180402 (2002)



Frequencies from “2nd order Bogoliubov”

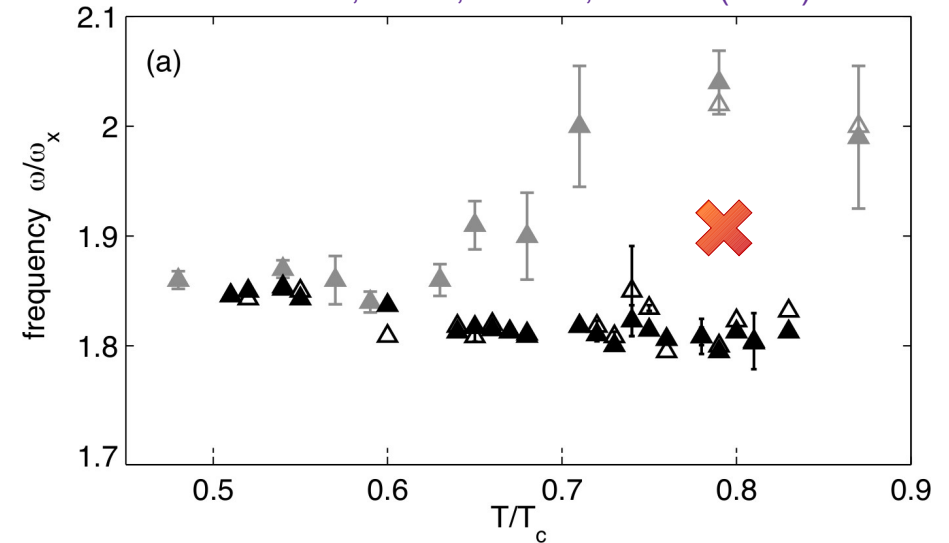
Morgan, Rusch, Hutchinson, Burnett PRL 91, 250403 (2003)



Flexible simulations

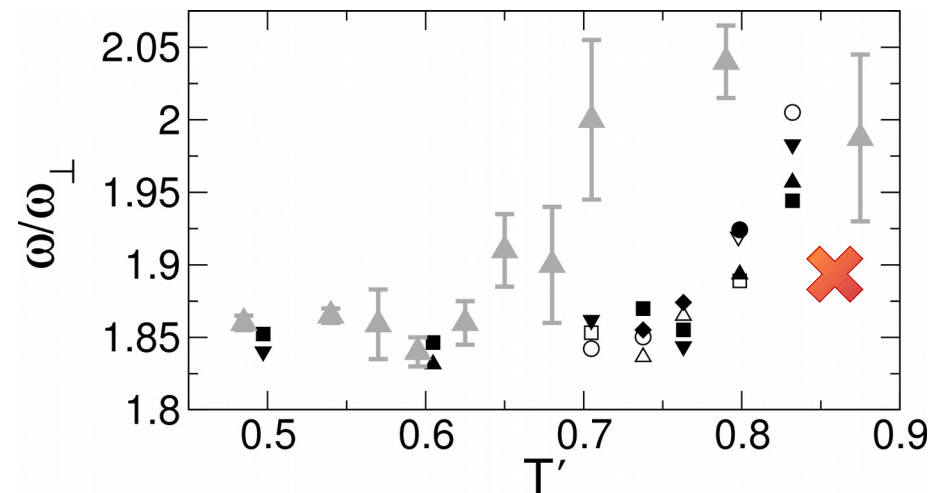
Classical fields + HF

Bezett, Blakie, PRA 79, 023602 (2009)

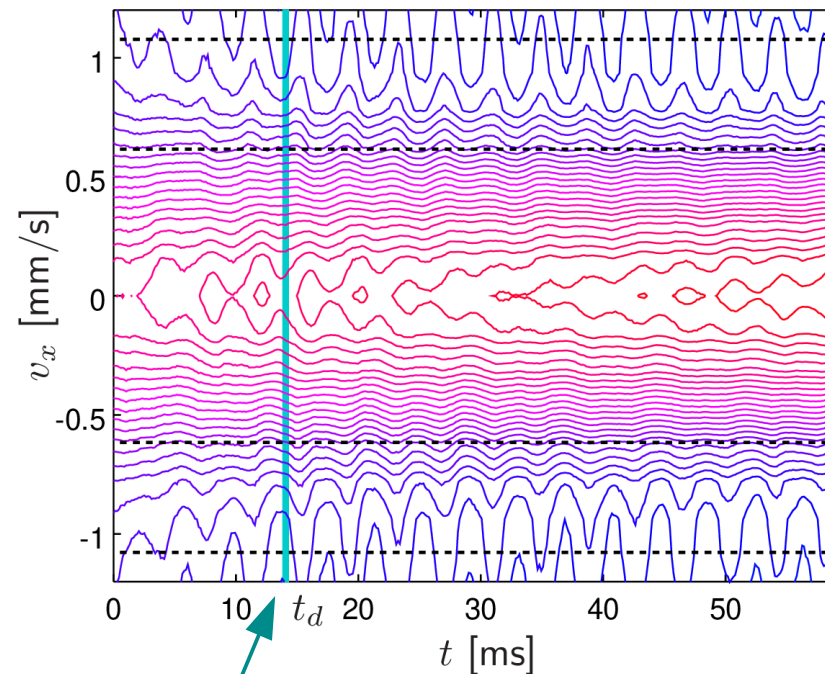
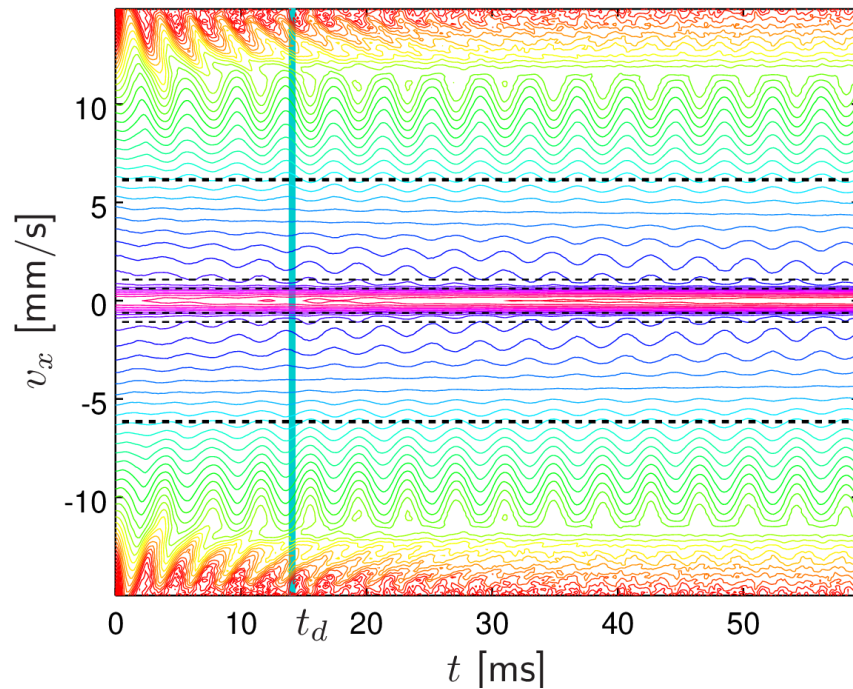


Classical fields (higher cutoff)

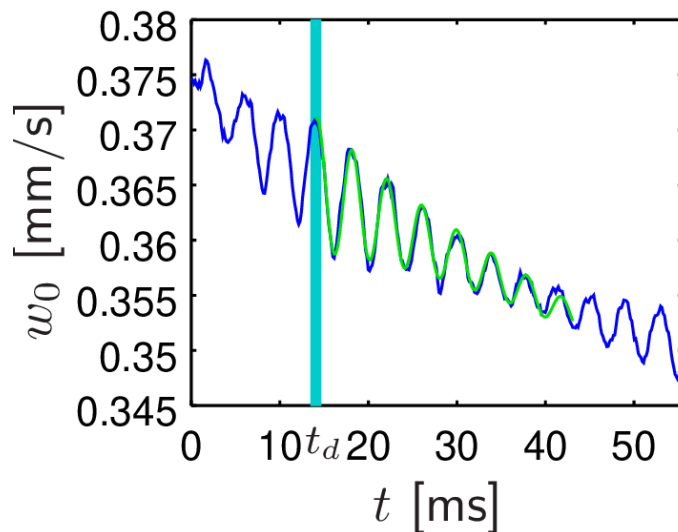
Karpiuk, Brewczyk, Gajda, Rzażewski, PRA 81, 013629 (2010)



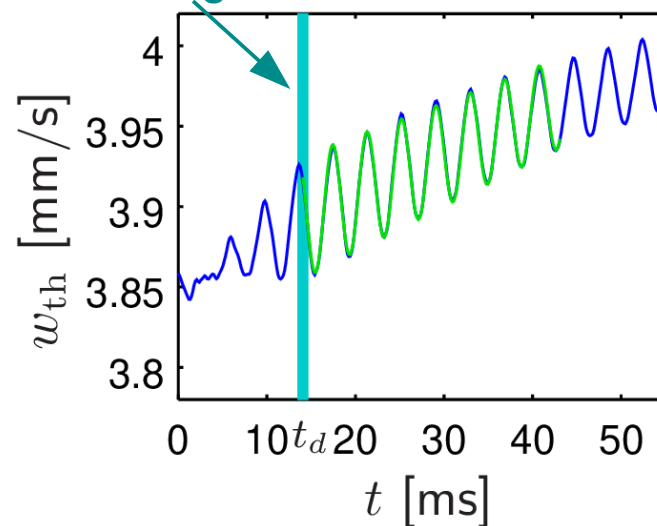
Contours of log density during evolution:



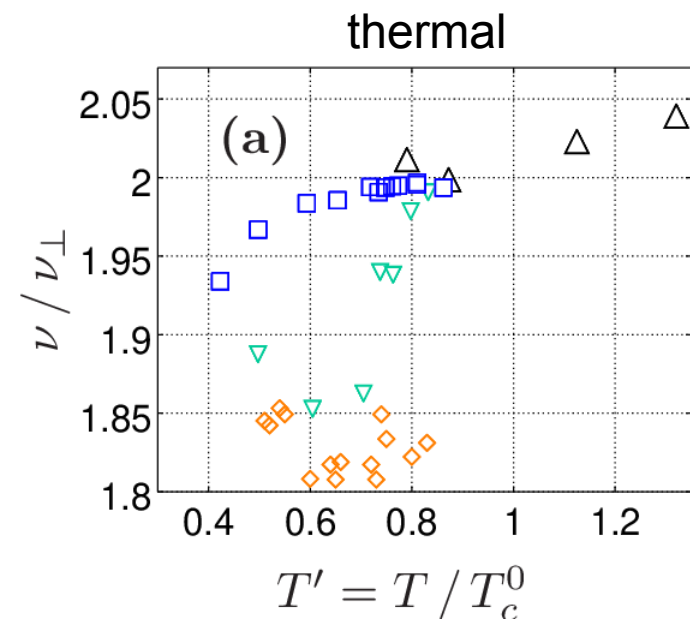
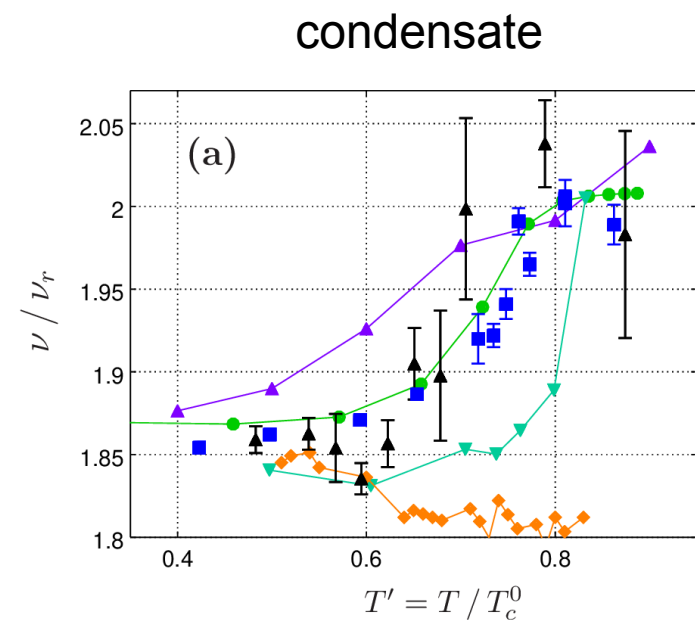
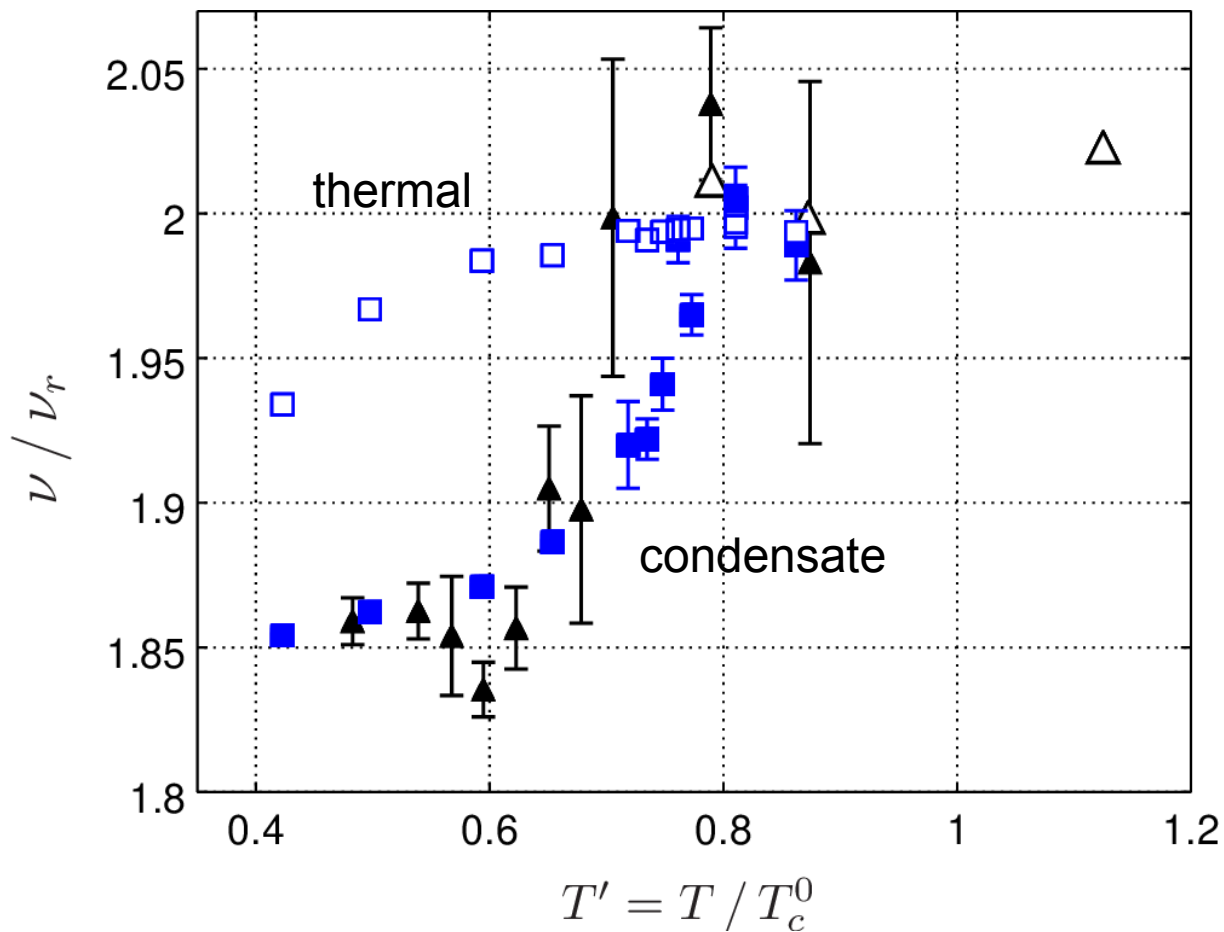
Fitting of width:



End of driving:



Results: $m=0$ mode frequency



Experiment: Jin, Matthews, Ensher, Wieman, Cornell, PRL 78, 764 (1997)

[This work \(2018\)](#)

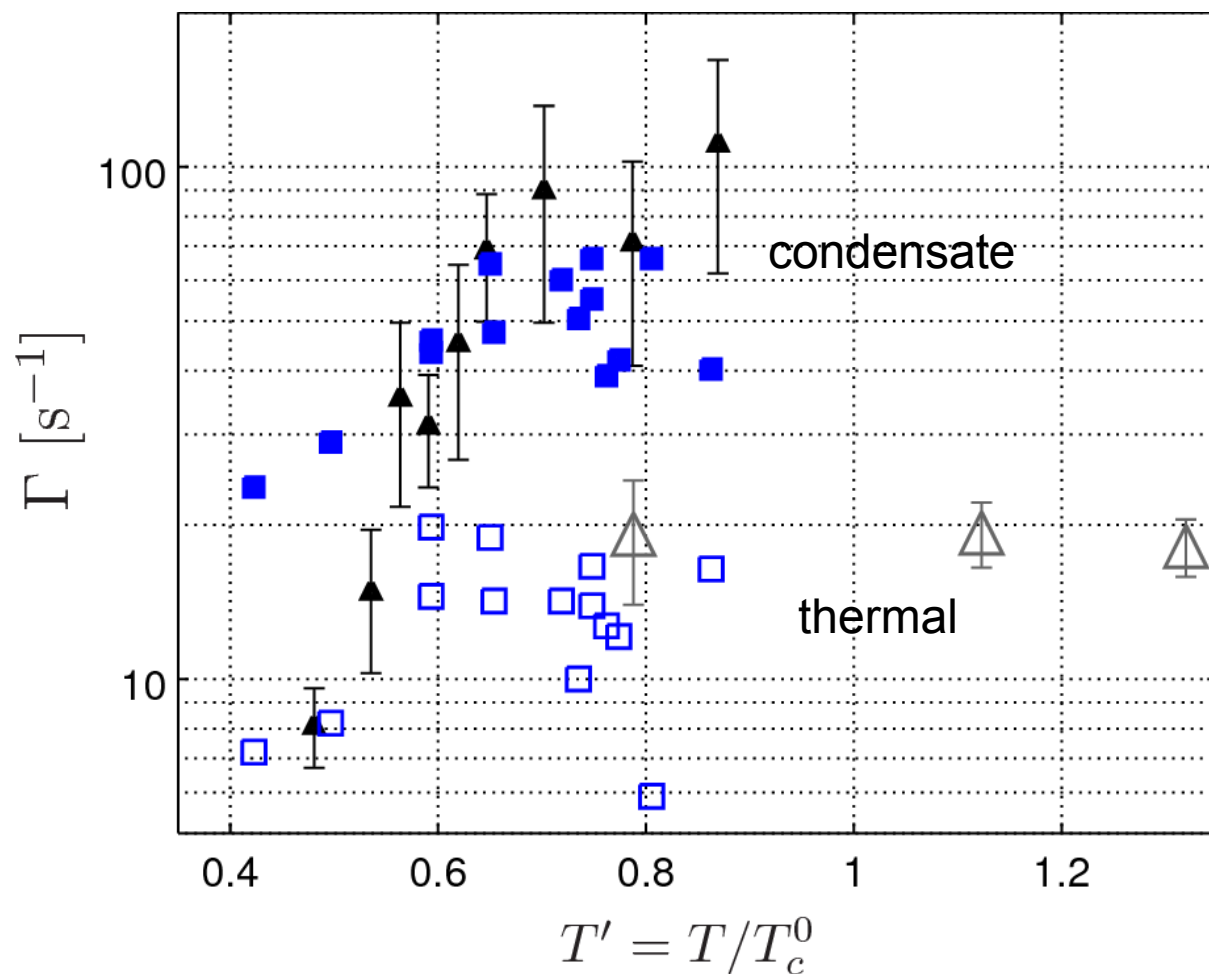
[ZNG: Jackson, Zaremba, PRL 88, 180402 \(2002\)](#)

[2nd order Bogoliubov: Morgan, Rusch, Hutchinson, Burnett PRL 91, 250403 \(2003\)](#)

[PGPE+HF: Bezettl, Blakie, PRA 79, 023602 \(2009\)](#)

[PGPE: Karpiuk, Brewczyk, Gajda, Rzażewski, PRA 81, 013629 \(2010\)](#)

$m=0$ mode damping



Experiment: Jin, Matthews, Ensher, Wieman, Cornell, PRL 78, 764 (1997)

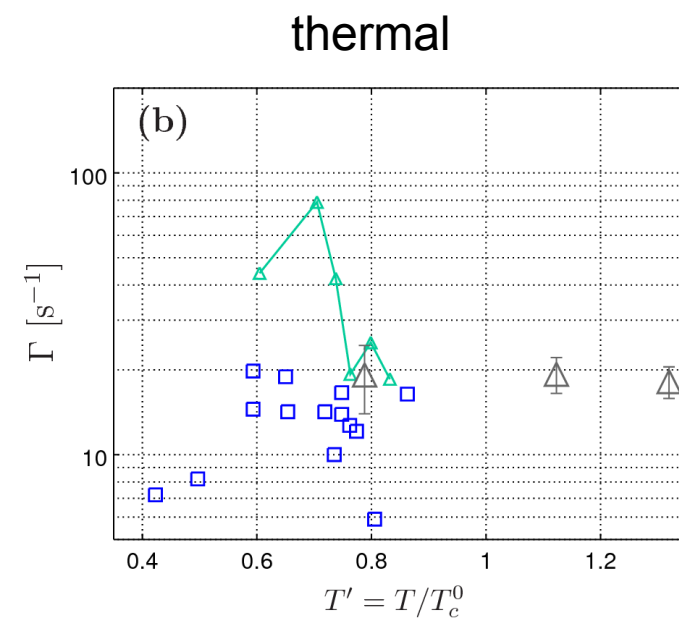
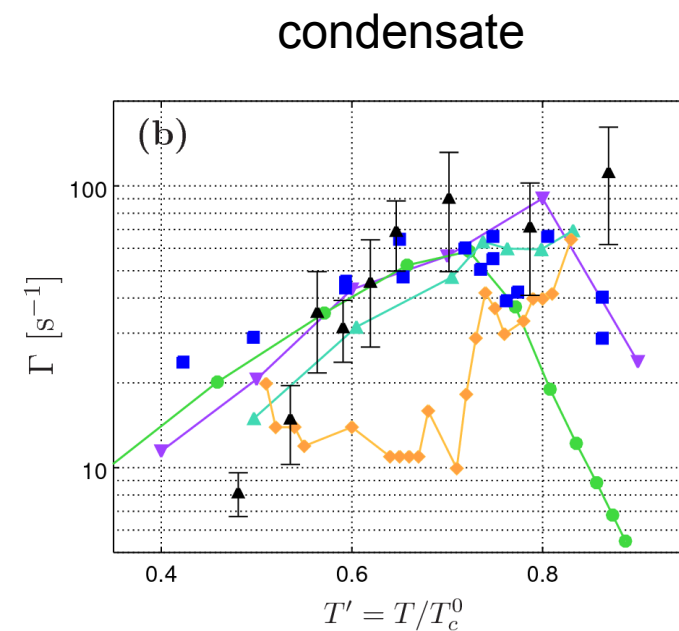
[This work \(2018\)](#)

[ZNG: Jackson, Zaremba, PRL 88, 180402 \(2002\)](#)

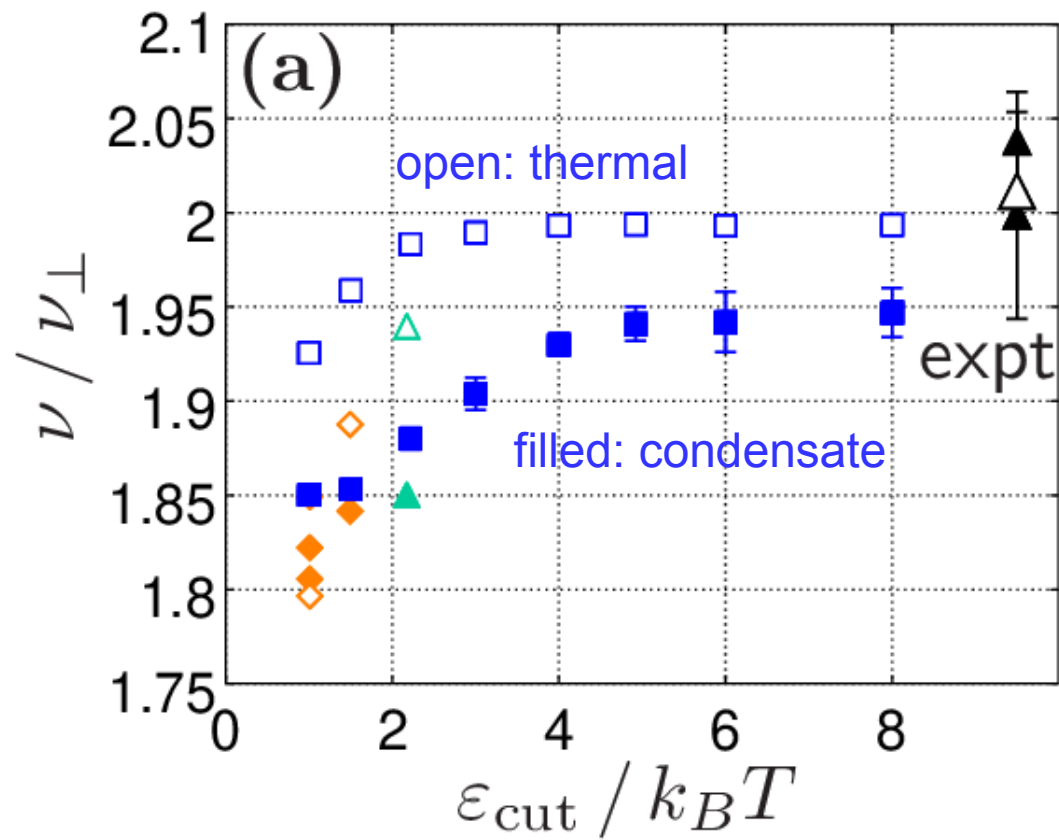
[2nd order Bogoliubov: Morgan, Rusch, Hutchinson, Burnett PRL 91, 250403 \(2003\)](#)

[PGPE+HF: Bezettl, Blakie, PRA 79, 023602 \(2009\)](#)

[PGPE: Karpiuk, Brewczyk, Gajda, Rzążewski, PRA 81, 013629 \(2010\)](#)



Cutoff dependence

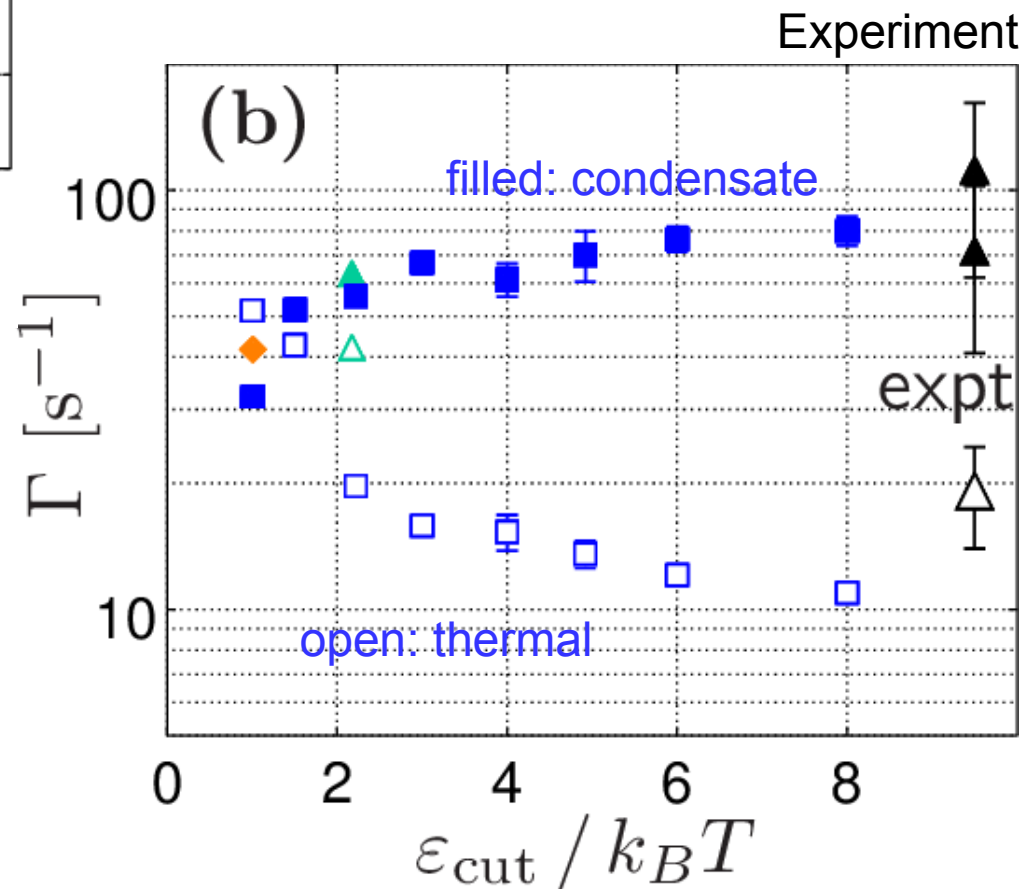


Experiment Jin *et al* 1997

Bezett + Blakie 2009

Karpiuk, Brewczyk, Gajda, Rzażewski 2010

This work



- Quantitative results with classical fields

Robust because no fitting of technical parameters such as cutoff

- Numerical effort comparable to SGPE

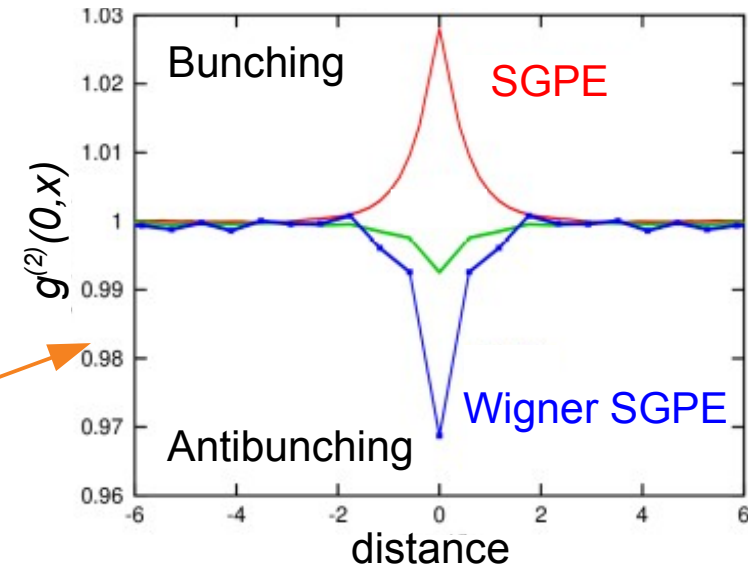
though, lattice may need to be larger

- Still lacks wave-particle duality

but this is harder to spot than expected

- NEXT: Wigner representation version

to include quantum fluctuations



Thanks to discussions with

Nick Proukakis
Mariusz Gajda
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Emilia Witkowska

Simon Gardiner
Crispin Gardiner
Matt Davis
Ashton Bradley
Blair Blakie

Krzysztof Gawryluk
Tomasz Karpiuk
Thomas Gasenzer
Andrew Daley
Krzysztof Pawłowski